Nothing But Coincidences

Adventures and Misadventure of Einstein's Point-Coincidence Argument

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point-coincidence argument (John Stachel)

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general covariance = no privileged coordinate systems

 Moritz Schlick (1917) found that Einstein point-coincidence remark aptly expressed the conceptual novelty of general relativity respect to previous spacetime theories.

Erich Kretschmann (1917) deemed the argument as trivial and actually valid in all spacetime theories

 Peter G. Bergmann (1960s) and the Syracuse Group made use of the expression coincidences, in a way that was much closer to Einstein's intentions.

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point-coincidence argument answer to the hole argument

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- The hole argument extremely successful
- The point-coincidence argument fundamentally neglected

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nothing but coincindeces = central message of general relativity

- $ds^{2} = \sum_{\mu\nu} g_{\mu\nu} dx_{\mu} dx_{\nu}$ metric field: device for extracting measurable distances ds between any two nearby points from their coordinate differences x_{ν}
- **gravitational field**: physical field that have certain values at a certain point with coordinates x_{ν} , just like the electromagnetic field.

- $ds^{z} = \sum_{\mu\nu} g_{\mu\nu} dx_{\mu} dx_{\nu}$ $\blacksquare \text{ metric field: device for extracting measurable distances } ds \text{ between any two nearby points from their coordinate differences } x_{\nu}$
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cartesian coordinates (x, y) on the map: $ds^2 = dx^2 + dy$

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Figure: cylindrical projection

$$ds^{2} = 4\left(1 + \frac{x^{2}}{R^{2}} + \frac{y^{2}}{R^{2}}\right)^{-2} dx^{2} + 4\left(1 + \frac{x^{2}}{R^{2}} + \frac{y^{2}}{R^{2}}\right)^{-2} dy^{2}$$

cartesian coordinates (x, y) on the map: $ds^2 = dx^2 + dy$



Figure: stereographic projection

$$ds^{2} = \left(1 - \frac{y^{2}}{R^{2}}\right)dx^{2} + \left(1 - \frac{y^{2}}{R^{2}}\right)^{-1}dy^{2}$$



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 $g_{\mu\nu} = map \ legend$

- \blacksquare the treasure is buried at x=45, y=85 in open ocean
- the numbers x, y are meaningless without the map legend $g_{\mu\nu}$

 $g_{\mu\nu}$ = map legend = treasure

• the treasure is the value of $g_{\mu\nu}$ at x = 45, y = 85 in open space

g_{$\mu\nu$} different values at the same point x = 45, y = 85 in different map legends

'values of the $g_{\mu\nu}$ at x, y' = meaningful statement?

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Early relativists thought in terms of geographical maps, rather than of the geometry of the surface

- **problem:** the Greenwich meridian intersects the equator at the point x = 45, y = 85 in one map but not in the other
- **solution:** x = 45, y = 85 is not the same physical point, the same physical point is where Greenwich meridian intersects the equator



intersections \implies coincidences

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intersections \implies coincidences

- classical mechanics \implies 'the same place in different times' meaningless
- special relativity \implies 'at the same time in different places' meaningless
- general relativity \implies at the same place at the same time meaningless



... rethinking the very notion of where something happens

Die Relativitäts-Theorie.¹)

Von

A. EINSTEIN in Prag.

" In mathematical physics, it is customary to relate things to coordinate systems [...] What is essential in this relating-to-something

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" dinate systems [...] What is essential in this relating-to-something is the following: when we state anything whatsoever about the lo-

Die Relativitäts-Theorie.¹)

Von

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" is the following: when we state anything whatsoever about the location of a point, we always indicate the coincidence of this point with some point of a specific other physical system. If, for exam-

Die Relativitäts-Theorie.¹)

Von

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" with some point of a specific other physical system. If, for example, I choose myself as this material point, and say, 'I am at this location in this hall,' then I have brought myself into spatial coincidence with a certain point of this hall, or rather, I have asserted this coincidence. This is done in mathematical physics by using three
Die Relativitäts-Theorie.¹)

Von

A. EINSTEIN in Prag.

" coincidence. This is done in mathematical physics by using three numbers, the so-called coordinates, to indicate with which points of the rigid system, called the coordinate system, the point whose location is to be described coincides.

"

In classical mechanics and special relativity 'where' something occurs were always meant as a physical coincidence with some material body of reference



(the surface of the Earth, the city of Zurich, or the walls of the hall in which Einstein was giving his lecture)

a good non-accelerating massive cubical scaffolding of rigidly connected measuring rods and clocks K(x, y, z, t)

- the readings of rods and clocks at rest with respect to K directly give the numbers x,y,z,t.
- the components of $\vec{E}(x, y, z, t)$ and $\vec{B}(x, y, z, t)$ with respect to K can be determined by observing the paths of test particles with respect to K.
- the paths of test particles are successive encunters with certain points of K when the handles of a clock are on a certain position on the dial.
- theoretical asymmetry between a good acceleration-free K(x, y, z, t)and the bad coordinate system $K'(\overline{x', y'}, z', t')$ accelerating with respect to the former

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Constitution One might wonder whether such pedantic physical definitions for the time and space coordinates are really necessary, i.e., whether it is really necessary to burden the beautiful and airy concepts of space and time with cumbersome rigid bodies and clocks. In my opinion it is not necessary, but it is advantageous to proceed in this way. One can in fact treat x, y, z as pale mathematical auxiliary quantities (parameters), that have meaning only because they facilitate the formulation of the physical laws [...] Anyway, I believe that this considerations and definitions about space and time are sufficient only in as much one forgoes the introduction of the gravitation in system of relativity theory

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- 1. x, y, z, t as mathematical parameters: expr. $\vec{E}(x, y, z, t) \rightarrow$ mere mathematical function that can't be experimentally established
- 2. x, y, z, t as readings on rods and clock: espr. $\vec{E}(x, y, z, t) \rightarrow$ has a physical meaning



the universal nature of gravitation made the interpretation (1) impossible.

It becomes impossible to establish whether K is a 'good' coordinate system in which nonaccelerated rods and clocks at rest reliably read coordinate differences or a 'bad' one K' accelerating in the opposite direction, in which coordinate numbers do not directly denote distances.

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 $ds^2 = \sum g_{\mu\nu} dx_{\mu} dx_{\nu}$ $\mu\nu$

Figure: Zurich Notebook

2 = E Gun day day (12) 18 XV age 91 agy x' B. 12 202 X3 age EE gu darden EEggo d'x' dx's = E E E Geo x nas day day Gen = E & Gooder ton analog Gen E & Goo Bed Am x = 5 x ro xo Inequalfall free dae Gue 2 . x $c^{2} = 2c \frac{2c}{2} \frac{2c}{2}$ do 2 = E Gen day days 4(c2) = 2 grudie + 20 00 grady a quely = grout(c2) = 2 c grad c 48 - 34

$$ds^2 = \sum_{\mu\nu} g_{\mu\nu} dx_{\mu} dx_{\nu}$$

Figure: Zurich Notebook



• metric field: extract measurable distances dsbetween any two nearby points from their coordinate differences dx_{ν}

 $ds^2 = \sum g_{\mu\nu} \, dx_\mu dx_\nu$

• gravitational field⁷: physical field that have certain values at a certain point with coordinates x_1, x_2, x_3, x_4 , just like the electromagnetic field.

 $g_{\mu\nu}(x_1, x_2, x_3, x_4)$



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...there cannot exist relationships between the spacetime coordinates x_1, x_2, x_3, x_4 and the results of measurements obtainable by means of measuring rods and clocks that would be as simple as those in the old relativity theory

"

labels \neq distances

Spannungs-Energie-Tensor des materiellen Vorganges

Den Tensor 0_e, nennen wir den (kontravrianten) Spannage-Energielensor der materiellen Strömung. Der Gleichung (10) schreiben wir einen Gültigkeitsbereich nu, der über den speziellen Fall (2) der Strömung inkohärente Massen weit hinzungelt. Die Gleichung stallt allgemein die Energieblanz zwischen dem Gravitationsfelde und einem belichigten materiellen Vorgang darz nur ist für 0_e, der dem jeweilen betrachieten materiellen System entsprechende Spannunge. Energitensore einzussten. Die erste Summe in der Gleichung ethalt die örtlichen Ableitungen der Spannungen bwr. Energiestrundichte und isreitlichen Ableitungen der Impulie bwr. Energiestrundichte gleich Summe ist ein Ausdruck für die Wirkungen, welche vom Schwerefelde auf den materiellen Vorgang übertragen werden.

§ 5. Die Differentialgleichungen des Gravitationsfeldes

Nachdem wir die Inpuls-Energiegieichung für die materiellen Vorginge (mechanische elektrische und andere Vorgränge) mit bezug auf das Gravitationsfeld aufgestellt haben, bleibt uns noch folgende Aufgabe. Es sei der Tensor Θ_{μ} , für den materiellen Vorgzag gegeben. Welches sind die Differentaligiebungen, wielbe die Größen Ω_{μ} , ch das Schwerefeld zu bestimmen gestatten? Wir suchen mit anderen Worten die Verallgemeinerung der Doissonsches Gliechung

$\varDelta \varphi = 4\pi k \varrho.$

Zur Lösung dieser Aufgabe haben wir keine so vollkommen zwangläufige Methode gefunden, wie für die Lösung des vorhin behandelten Problems. Es war nötig, einige Annahmen einzuführen, deren Richtigkeit zwar plausibel erscheint, aber doch nicht erident ist.

Die gesuchte Verallgemeinerung wird wohl von der Form sein

(11)
$$\mathbf{x} \cdot \Theta_{\mu\nu} = \Gamma_{\mu\nu}$$
,

wo z eine Konstante, $I_{\gamma_{e}}$ ein konstravarianter Tensor zweiten Ranges ist, der durch Differentialoperstoom aus dem Frankannenklassor $J_{\gamma_{e}}$ herorogett. Dem Newton-Poissonnehen Gesetz entsprechend wird mang senist sein im fordera, daß disse fleichtungen (11) zweiter Orf- $I_{\gamma_{e}}$ en finden, der eine Veraligemeierung von $J_{\gamma_{e}}$ at finden, der eine Veraligemeierung von $J_{\gamma_{e}}$ at finden, der eine Veraligemeierung von $J_{\gamma_{e}}$ at disch die die die Australie die Veralie die Veralie die Geschleichtung Hann allerdingen Transformationen gegenüber als Tensor erweist.¹ A priori kun allerdingen beicht in Abrede gestellt werden, abgichkeit, daß die sein Konnten. En besteht daher immer noch die Möglichkeit, daß die

1) Vgl. II. Teil, § 4, Nr. 2.

field equations relating the source variables (charge density, mass energy distribution, etc.) to the field variables (gravitational field, electromagnetic field, etc.)

$$\Delta \varphi = 4\pi k \varrho$$

$$\kappa \Theta_{\mu\nu} = \Gamma_{\mu\nu}$$

- I want to know the the values of g_{µν} at a point x₁, x₂, x₃, x₄
- The $g_{\mu\nu}$ tell me where the point x_1, x_2, x_3, x_4



Entwurf-theory field equations of limited covariance \rightarrow Mercury's anomalous perihelion precession

Is the static gravitational field $g_{\mu\nu} =$ 1,1 to 3, $g_{44} =$ f(x, y, z) a particular solution? Or is it the general solution expressed in particular coordinates?^a

[Besso:] The requirement of [general] covariance of the gravitational equations under arbitrary transformations cannot be imposed: if all matter [is given] were contained in one part of space and for this part of space a coordinate system [is given], then, outside of it. the coordinate system, except for boundary conditions be chosen arbitrarily, so that a unique determinability of the g's cannot be obtained

Die Aufordering In Allymaine Covarian In Geowartions gleichen für. beliden Tronsformation an doner wicht auf, geschelt weden: When in eren Raunes alle ustur gyption ist and fin Insur Kore Coordination System, 20 Rounde trok aussichalt Disselben Das Coordinations form work in wesentlother abjeschen von by Preus bedizungen, beliching gun Talt werden, for Such Doi g belieting educe to loss com eviduting Destinue backet du go witht cole ficture house. how bit of widing bestoment and, condim un bit in growhations roum berbachtbacen Erscheinungen z. B. Der Beurg Des austenstellen Preuchtes, entissen is dem High wichter Denen Such De cher Bereg est auch che cher Beweger Will gageben. 350 im Detrin und vergen eine Lisury He so ist doesn's selbe Sebilde and ein Losury in 2, The ; Ke aba and cire L'suy in 1. Est

Die Aufordener In allymeinen Covariary der Georstotionsyleicher für belid oge Tronof mation en former wicht auf, gebilt werden: Uhen in einem Aannes alle usture gyptim ist and fin Dusin Tellon. Conditionten System, 20 Rounde Inch aussichalt Disselben Das Coordinatursystem node, in wescuttother abjeschen von Sun freux bedizungen, belerking gun Talt merten, we druck de q beliching estre to loss com endutig Destimu backet der g " wicht cole freten kounte. how bit of widing bestorent ever condin un bit in growtations roum berbachthours Ersteinungen, g. O. Jor Dearry Des such and the Puelkes, unter a sen fritt wilds "Sun Duck In Bereger Toll gegeben. Jet in Dervicer in inter iget 1 eine Lisury He so ist doesn's selbe Sebilde ouch eine Losny in 2, 3/2; Ke aba auch cire L'suy in 1. Est

[Besso:] It is, however, not necessary that the themselves are determined uniquely, only the observable phenomena in the gravitation space, e.g., the motion of a material point, must be

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[Einstein:] Of no use, since with a solution a motion is also fully given. If in coordinate system 1, there is a solution K_1 , then this same construct is also a solution in 2, K_2 ; K_2 however, also a solution in 1 2) In den letzten Tagen fand ich den Beweis datür, daß eine derartige allgemein kovariante Lösung überhaupt nicht existieren kann.

herab. Das ganze Problem der Gravitation wäre also befriedigend gelöst, wenn es auch gelänge, bezüglich beliebiger Substitutionen kovariante Gleichungen zu finden, welchen die das Gravitationsfeld selbst bestimmenden Größen g_{av} genügen. In dieser Weise gelang es uns aber nicht, das Problem zu lösen³). Die Lösung gelang aber in der Weise, daß nachträglich wieder das Bezugssysten spezialisiert wurde. Zu diesem Wege gelangt man ungezwungen durch folgende Überlegung. Es ist klar, daß für irgend-

1) Kottler. Über die Raumzeitlinien der Minkowskischen Welt. Wien. Ber. 121, 1912.

2) In den letzten Tagen fand ich den Beweis dafür, daß eine derartige allgemein kovariante Lösung überhaupt nicht existieren kann. In the last days found a proof that such a generally covariant solution to the problem cannot exist at all

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Es gebe in unserer vierdimensionalen Mannigfaltigkeit einen Teil	Einstein inverted		
also die Q verschwinden Durch die außerhelb L gegebenen Q sind	Posso's argument		
and the $\Theta_{\mu\nu}$ verschwindten. Dutch die auschnate D gegebenen $\Theta_{\mu\nu}$ sind gemäß ungerer Annahme übereil also auch im Innern von L die w	besso's argument		
yollkommen bestimmt. Wir denken uns nun statt der ursprünglichen			
Koordinaten x neue Koordinaten x' eingeführt von folgender Art.	_		
Außerhalb L sei überall $x_i = x'_i$; innerhalb L aber sei wenigstens für			
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wenigstens für einen Teil von $L \gamma'_{\mu r} + \gamma_{\mu r}$ ist. Andererseits ist überall			
$\Theta'_{\mu\nu} = \Theta_{\mu\nu}$, nämlich außerhalb L , weil für dieses Gebiet $x'_{\nu} = x_{\nu}$ ist, —			
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einzuschranken. Diese Einschrankung wird in unserer Arbeit dadurch			
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Ans dissel Es gebe in unserer vierdimensionalen Ma	nnigfaltigkeit einen Teil		
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		T T Y			

a) Wenn das Bezugssystem ganz willkürlich gewählt wird, dann können die g" durch die \mathbb{Z}_{n} überhaupt nicht vollständig bestimmt sein. Man denke sich nämlich die \mathbb{X}_{n} und g" überall gegeben, und es mögen in einem Teil Φ des vierdimensionalen Raumes alle \mathbb{Z}_{n} , und g", überall gegeben, und ein neues Bezugssystem einführen, welches außerhalb Φ mit dem ursprünglichen vollkommen übereinstimmt, innerhalb Φ aber (ohne Ver letzung der Stetigkeit) von ihm verschieder ist. Bezieht man nun alles auf dieses neue Bezugssystem, wobei die Materie durch \mathbb{Z}_{n} , das Gravitationsfeld durch g" ausgedrückt wird so ist zwar überall

dagegen werden im Innern von Φ die Gleichungen

$$g_{\mu\nu} = g_{\mu\nu}$$

sicherlich nicht alle erfüllt sein¹). Hieraus folgt die Behauptung.

Will man erzielen. daß eine vollständige Bestimmung der $g_{\mu\nu}$ (Gravitationsfeld) aus den $\mathfrak{T}_{\mu\nu}$ (Materie) möglich sei, so kann dies nur durch Beschränkung in der Wahl des Bezugssystems erreicht werden.

I) Die Gleichungen sind so zu verstehen, daß auf den linken Seiten jeweilen den unabhängigen Variabeln x, dieselben Zahlenwerte erteilt werden wie auf den rechten Seiten den Variabeln x.

$$T'_{\mu\nu} = T_{\mu\nu}$$



I) Die Gleichungen sind so zu verstehen, daß auf a) Wenn daden linken Seiten jeweilen den unabhängigen Variabeln x'_{j} kürlich gewäh g_{un} durch die ständig bestin Seiten den Variabeln x_{j} .

nämlich die \mathfrak{L}_{p}^{-} und g_{s}^{-} uortan gegeven, unves es mögen in einem Teil Φ des vierdimensionaler Raumes alle \mathfrak{T}_{s}^{-} verschwinden. Ich kann nur ein neues Bezugssystem einführen, welches außerhalb Φ mit dem ursprünglichen volkkommen übereinstimmt, innerhalb Φ aber (ohne Ver letzung der Steiigkeit) von ihm verschieder ist. Bezieht man nun alles auf dieses neue, Bezugssystem, wobei die Materie durch \mathfrak{T}_{s}^{-} das Gravitationsfeld durch g_{ss}^{-} ausgedrückt wird so ist zwar überall

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The §12 Argument

§ 12. Beweis von der Notwendigkeit einer Einschränkung der Koordinatenwahl.

Wir betrachten einen endlichen Teil Z des Kontinuums, in welchem ein materieller Vorgang nicht stattfindet. Das physikalische Geschechen in X ist dann vollständig bestimmt, wenn in bezug auf ein zur Beschreibung benutztes Koordinatensystem K die Größen g_{ω} als Funktion der z. geschen werden. Die Gesamtheit dieser Funktionen werde symbolisch durch G(z) bezeichnet.

Es werde ein neues Koordinatensystem K' eingeführt, welches sußerhalb zmit K übereinstimme, innerhalb z barv ron K abweiche, derart, daß die auf K' bezogenen g'_* wie die g'_* (nebst ihren Ableitungen) überall stetig sind. Die Gesamtheit der g'_* bezeichnen wir symbolisch durch G'(z). G'(z) und G(z) beschreiben das nämliche Gravitationafeld. Ersetzen wir in den Funktionen g'_* die Koordinaten z'_* durch die Koordinaten z_* . d. b. bilden wir G'(z), so beschreibt G'(z) ebenfülls ein Gravitationafeld bezüglich K_* welches aber nicht übereinstimmt mit dem tatschlichen (Dzw. ursprünglich gegebenen) Gravitationsfelde.

Setten wir nun voraus, daß die Differentialgleichungen des Gravitationsföldes allgemein kovräntat sind, so sind sie für G'(z) erfüllt (bezüglich X^{*}), wenn sie bezüglich K für G'(z) erfüllt. Bezüglich K existierten dann die voneinander verschiedenen Lösungen G(z) und G'(z), totzdem an den Gebietzgenzen beide Lösungen G(z) und G'(z), totzdem an den Gebietzgenzen beide Lösungen til diereinstimmten, d. h. durch allgemein kovariante Differentialgleichungen für das Gravitationafeld kann das Geschehen in demselben nicht eindeutig festgelegt werden.

Verlangen wir daher, daß der Ablauf des Geschehens im Giravitationsfelde durch die auftrautellenden Gesetze volltstäußig bestimmt sei, so sind wir genötigt, die Wahl des Koordinatensystems derart einzuschränken, daß es ohne Verletzung der einschränkenden Bedingungen numöglich ist, ein neues Koordinatensystem K' von der vorhin charakterisierten Art einzuführen. Die Fortsetzung des Koordinatensystems ims Innere eines Gebietes 2 hinein dart nicht willkrüch sein. region Σ of spacetime

$$G(x) \to G'(x)' \to G'(x)$$

C There are then two different solutions G(x) and G'(x) relative to K, even though the solutions coincide on the boundary of the domain Σ

Back to General Covariance

Zur allgemeinen Relativitätstheorie.

Von A. Einstein.

In den letzten Jahren war ich bemüht, auf die Voraussetzung der Relativität auch nicht gleichförmigen Bewegungen eine allgemeine Realivitätsheorie zu gründen. Ich glaubte in der Tat, das einzige Gravitationsgesetz gefunden zu haben, das dem sinngemäß gefaßten allgemeinen Relativitätspositulate entspricht, um dusehte die Notwendigkeit gerade dieser Lösung in einer im vorigen Jahre in diesen Sitzungsberichten erschienena Arbeit¹ durzutun.

Eine erneute Kritik zeigte mir, daß sich jene Notwendigkeit auf dem dort eingeschlagenen Wege absolut nicht version 1084; daß dies doch der Fall m. sein schlen, beruhte auf Irrtum. Das Postulat der Relativitä, sowei ich es dort gefordert habe, ist stes erfüllt, wenn man das Hasurtossche Prinzip zugrunde legt; es liefert aber im Wahrheit keine Handhabe frei eine Ermittelung der Haurtosschen Funktion H des Gravitationsfeldes. In der Tat dreckt die die Wahl von H einschränkende Gielehung (77) a. O. nichts anderes aus, als daß H eine Invariante beträglich linearer Transformationen sein solj, welche Forderung mit der der Relativität der Beschennigung nichts zu schaffen hat. Ferner wird die durch Gielchung (75) a. O. getroffene Wahl durch Gielchung (75) keinswege fretzelet.

Aus diesen Gründen verlor ich das Vertrauen zu den vom mir aufgestellten Pödgleichungen vollständig um dusuche nach einem Wege, der die Möglichkeiten in einer natürlichen Weise einschräufst. So gelangte ich zu der Forderung einer allgemeineren Kovarianz der Feldgleichungen zurück, von der ich vor drei Jahren, als ich zusammen mit meinem Preunde Grossanzs arbeitete, nur mit schwerem Herzen abgegangen war. In der Tat waren wir damals der im nachfolgenden gegebenen Lösung des Problems bereits ganz nach gekonderne.

Wie die spezielle Relativitätstheorie auf das Postulat gegründet ist, daß ihre Gleichungen bezüglich linearer, orthogonaler Transforfour papers in November 1915 ...

¹ Die formale Grundlage der allgemeinen Relativitätstheorie. Sitzungsberichte XLI, 1914, S. 1066-1077. Im fölgenden werden Gleichungen dieser Abhandlungen beim Zitieren durch den Zustat: -a. a. -v no solchen der vorliegenden Arbeit unterschieden.

Back to General Covariance

"

A point mass, the sun, is located at the origin of the coordinate system. The gravitational field this point mass produces can be calculated from the [field equations] [...] Nevertheless, we should consider that the $g_{\mu\nu}$ are still not completely determined mathematically by [the field] equations [...] Yet we are justified in assuming that all these solutions can be reduced to one another by such transformations that they are distinguished (by the given boundary conditions) formally but not, however, physically, from one another

Einstein, 1915

Back to General Covariance

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Einstein, 1915

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Einstein, 1915



In 1914 Einstein developed an almost philosophical proof [against general covariance] (I mean e.g. §12 of his paper of 19.XI.1914. Is this proof correct? Ehreniest to Lorentz, Dec. 23, 1915

C What is observed here if we neglect, at first, all direct metrical evaluations — is only the complete or partial spatiotemporal coincidence [Zusammenfallen] or noncoincidence [Nichtzusammenfallen] of parts of the measuring instrument with parts of the measured object Kretschammn1915

12/21/1915

Deckung gebrachten Teile des Meßgegenstandes. Beobachtet wird hierbei - wenn wir zunächst von allen direkten Größenschätzungen absehen - nur das völlig oder teilweise erreichte räumlich-zeitliche Zusammenfallen oder Nichtzusammenfallen von Teilen des Meßinstrumentes mit Teilen des Meßgegenstandes. Oder allgemein: rein topologische Beziehungen zwischen

Erich Kretschman

Moritz Schlig

All measurements happen in a way that spatial coincidences [Coinzidenzen] (Galvanometer deflections, positions on a clock dial etc.) can be observed ... (Zurich 1910)

"



- **(** In §12 of my paper of last year, everything is correct [...] A contradiction to the uniqueness of the event does not follow at all from the fact that both systems G(x) and G'(x), related to the same frame of reference, satisfy the conditions of the grav. field. [...]
 - the reference system has no real meaning
 - 2. that the (simultaneous) materialization of two different g systems [...] within the same area of the continuum is [...] impossible

In place of §12 the following consideration must appear. Whatever is physically real in events in the universe (as opposed to that which is dependent on the choice of a reference system) consist in spatio-temporal coincidences^{*}

*and in nothing else!

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In place of §12 the following consideration must appear. Whatever is physically real in events in the universe (as opposed to that which is dependent on the choice of a reference system) consist in spatiotemporal coincidences⁴ [...] [Two inter-transformable $g_{\mu\nu}$ -systems] are entirely equivalent

*and in nothing else!

"

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In place of §12 the following consideration must appear. Whatever is physically real in events in the universe (as opposed to that which is dependent on the choice of a reference system) consist in spatiotemporal coincidences^{*} This is because they have in common all the spatial-temporal point coincidences, that is, all the observables

*and in nothing else!

"

C I will defend the philosophy of §12 against your refutation

"

Ehrenfest to Einstein, Jan. 1, 1916

star emitting infinitely weak light



I will defend the philosophy of §12 against your refutation

"

Ehrenfest to Einstein, Jan. 1, 1916

light waves in <u>empty</u> <u>space</u>



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Ehrenfest to Einstein, Jan. 1, 1916

an aperture, two telescopes a photographic plate



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Ehrenfest to Einstein, Jan. 1, 1916

The situation described using $x_1 \dots x_4, T_{\mu\nu}, g_{\mu\nu},$ etc. The situation described $x_1 \dots x_4, T_{\mu\nu}, g_{\mu\nu}, g_{\mu\nu},$ etc.

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Ehrenfest to Einstein, Jan. 1, 1916

The situation described using $x_1 \dots x_4, T_{\mu\nu}, g_{\mu\nu}$, etc.



G Bravo the photrogr. plate is darkened (a true 'coincidence explanation').

I will defend the philosophy of §12 against your refutation

"

Ehrenfest to Einstein, Jan. 1, 1916

The situation described using $x_1 \dots x_4, T_{\mu\nu}, g_{\mu\nu}$, etc.



Then let's see the philosophy of §12 at work $G(x) \to G'(x') \to G'(x)$

I will defend the philosophy of §12 against your refutation

"

Ehrenfest to Einstein, Jan. 1, 1916

In the cross-hatched region occupied by matter, the coordinate system is kept fix



C Then let's see the philosophy of §12 at work $G(x) \to G'(x') \to G'(x)$

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"

Ehrenfest to Einstein, Jan. 1, 1916

whereas in the noncross-hatched empty space the latter is strongly changed



C Then let's see the philosophy of §12 at work $G(x) \to G'(x') \to G'(x)$

You look at me laughing quietly and you say 'go ahead, young friend, and describe if you like, the empty space

with old coordinates xand brand new G'(x)

nothing observable, no 'coincidence' would change!? Now I'm astonished and *angry* over your laugh and

I claim with clenched fist:

'If with the old x and the old G(x) and the new G'(x) one calculates the darkening of the plate, then one should calculate the nondarkening of the plate with old x and new G'(x)'

"

Ehrenfest to Einstein, Jan. 1, 1916

C The root of your difficulty lies in the fact that you instinctively treat the reference system as something real

Einstein to Ehrenfest, Jan. 5, 1916

- 4times

 draw star, the aperture and the plate onto completely deformable tracing paper



C The root of your difficulty lies in the fact that you instinctively treat the reference system as something real

Einstein to Ehrenfest, Jan. 5, 1916



 make another tracing on the letter paper



^athe coordinates of the star, the aperture and the plate are the same, as well use the boundary conditions at infinity

C The root of your difficulty lies in the fact that you instinctively treat the reference system as something real

))

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C The root of your difficulty lies in the fact that you instinctively treat the reference system as something real

Einstein to Ehrenfest, Jan. 5, 1916



34/51

When you relate the figure once again to the orthogonal writing paper coordinates, the solution is mathematically different from the original, and naturally also with respect to the $g_{\mu\nu}$. But physically it is exactly the same, since the writing paper coordinate system is only something imaginary. Always the same points are illuminated on the plate. [...] As long as the drawing paper, i.e., 'the space', is unreal, both diagrams do not differ at all. Always the same points are illuminated on the plate [...] It all depends on coincidences, e.g., whether the plate points are hit by the light or not

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- The same diagram's lines do not intersect at the same point of the latter paper.
- The same point is there where the same diagram worldlines intersect.

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Private PCA: The Lorentz-Ehrenfest Correspondence



Lorentz-Ehrenfest Correspondence, January 1916^{*}

[†]Thanks to A.J. Cox!.

Private PCA: The Lorentz-Ehrenfest Correspondence

'matter-free' field:

$$g_{\mu\nu} = F(x_{\alpha}) \text{ (field I)}$$

$$g'_{\mu\nu} = F'(x'_{\alpha})$$

$$F'(x_{\alpha}) \text{ (field II)}$$

This is a new solution, differing from the first. [...] (I) and (II) now in fact differ physically, since in field (I) a material point moves uniformly along a straight line, whereas one can easily see that this is not the ease in field (II). [...] Of course there exists a similar indeterminacy also in other cases, e.g. the sun's field, calculated by Einstein.

Lorentz to Ehrenfest, Jan. 9, 1916
x_1, x_2, x_3, x_4





 x_1, x_2, x_3, x_4





 x_1, x_2, x_3, x_4

$$ds = 0$$

- light rays traverse different
 - x_1, x_2, x_3, x_4 values in empty space
- light rays meet the same point x_1, x_2, x_3, x_4 on the plate in matter region

+Stene - tes Stuck x_1, x_2, x_3, x_4

$$ds = 0$$

 light rays traverse different

 x_1, x_2, x_3, x_4 values in empty space

light rays meet the same point x₁, x₂, x₃, x₄ on the plate in matter region

C Einstein is fully correct when he says: *A* and *B* produce the same darkening of the photographic plate

"

C I admit that we observe only 'coincidences'. [...] In connection with this I now realize that in the two fields I and II that I spoke of in my last letter, it is true that the separate phenomena do not take the same course in relation to the coordinate-system, but that in both of them the coincidences do occur in the same way. I was too much a prisoner of the idea that our equations must fully reproduce the relations between the phenomena and the chosen coordinate-system, whereas we can be happy if they duly reproduce the mutual relations between the phenomena.

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Lorentz to Ehrenfest, Jan. 10, 1916

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Lorentz to Ehrenfest, Jan. 10, 1916

- The claim that a planet is at certain point at a certain instant respect to the sun means that the astronomical sight lines up the planet and a fixed star.
- the trajectory of a light ray joining the star and the orbit of the planet, passing through a telescope co-moving with earth, and living marks on a photographic plate is interrupted by the passage of Venus when the handles of a clock is on a certain position on the dial.
- any statement about a position of the planet is ultimately a statement about the intersections among the worldline of the light ray coming from the star, the worldline of the planet, and that of the telescope, etc..
- statement about the intersection of worldlines reduce to statement about coincidences

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'Jupiter will be at φ, ϑ, r at the time t.'

- The claim that a planet is at certain point at a certain instant respect to the sun means that the astronomical sight lines up the planet and a fixed star.
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 statement about the intersection of worldlines reduce to statement about coincidences

- The question whether these worldlines will intersect at the same spacetime location φ, ϑ, r, t according to all inter-trasfromable $g_{\mu\nu}$ is meaningless
- the same spacetime location is where the same worldline intersect, which does not depend on the $g_{\mu\nu}$ used

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Private PCA: The Einstein-Lorentz Correspondence



Figure: ... Prachtswinkel auf diesem öden Planeten

I see that you have thought over the theory entirely and have familiarized yourself with the idea that all of our experiences in physics refer to coincidences

Einstein to Lorentz, Jan. 17, 1916

Physics. - "On Einstein's Theory of gravitation." 1. By Prof. H. A. LOBENTZ.

(Communicated in the meeting of February 26, 1916).

Now EINSTEIN has made the striking remark³) that the only thing we can learn from our observations and with which our theories are essentially concerned, is the existence of these coincidences.

⁸) In a correspondence I had with him.





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1916.

№ 7.

ANNALEN DER PHYSIK. VIERTE FOLGE. BAND 49.

1. Die Grundlage der allgemeinen Relativitätstheorie; von A. Einstein.

That this requirement of general co-variance, which takes away " from space and time the last remnant of physical materiality [Gegenständlichkeit], is a natural one, will be seen from the following reflexion. All our spacetime assessments [Konstatierungen]

"

" lowing reflexion. All our spacetime assessments [Konstatierungen] invariably amount to a determination of spacetime coincidences.

"

" If, for example, events consisted merely in the motion of material points, then ultimately nothing would be observable but the meetings [Begegnungen] of two or more of these points. More-

77

" meetings [Begeanungen] of two or more of these points. Moreover, the results of our measurings are nothing but assessments [Konstatierungen] of such meetings of the material points of our measuring instruments with other material points, coincidences between the hands of a clock and points on the clock dial

- coordinates x₁, x₂, x₃, x₄ do not tell us where such coincidences happen with respect to a given coordinate system;
- the four numbers x_1, x_2, x_3, x_4 are only a bookkeeping system to keep track of which coincidences we are referring to in a given coordinate system.

When we were describing the motion of a material point relative to a body of reference, we stated nothing more than the encounters of this point with particular points of the reference-body. [...] in conjunction with the observation of the encounter of the hands of clocks with particular points on the dials. [...] The following statements hold generally: Every physical description [...] refers to the spacetime coincidence.

Einstein, 1917, 1916

pre-GR: the material point coincide with a point of a scaffolding

When we were describing the motion of a material point relative to a body of reference, we stated nothing more than the <u>encoun-</u> ters of this point with particular points of the reference-body. [...] in conjunction with the observation of the encounter of the hands of clocks with particular points on the dials. [...] The following statements hold generally: Every physical description [...] refers to the spacetime <u>coincidence</u>.

Einstein, 1917, 1916

- pre-GR: the material point coincide with a point of a scaffolding
- GR: coincidence of trajectories of two points

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Einstein, 1917, 1916

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Reception of The PCA: Schlick and Kretschmann



- Moritz Schlick: the PCA is essential → the choice between different spacetimes which agree on point-coincidences is conventional^a
- Erich Kretschmann: the PCA argument is trivial \rightarrow the choince of the coordinate system to describe the same spacetime is conventional since all coordinate system agree on point-coincidences^b

^aSchlick, 1917. ^bKretschmann, 1917.



flapping rod paradox

- a rod at rest in one coordinate system.
- If one uses light rays as a standard, the rod is appears to be straight.
- introducing a fast changing gravitational field, the rod would appear flapping.



The two coordinate systems are indeed geometrically equivalent, however, physically, the first coordinate system is clearly better than the second one

I do not agree at all with your reflection about the bent (flapping) rod. All physical descriptions that yield the same observable relations (coincidences) are equivalent in principle, provided that both descriptions are also based on the same laws of nature. The choice of coordinates can have great practical importance from the point of view of clarity of description; in principle, though, it is entirely insignificant. It means nothing that 'arbitrary gravitational fields' occur, depending on the coordinate choice; [...] in principle, one can actually only learn something about the latter by eliminating the coordinates. The ghost of absolute space haunts your rod example.

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918.

No 4.

ANNALEN DER PHYSII a) Relativitätsprinzip: Die Naturgesetze sind nur Aussagen VIEBTE FOLGE BAND 55. er zeiträumliche Koinzidenzen; sie finden deshalb ihren einzig türlichen Ausdruck in allgemein kovarianten Gleichungen.

Prinzipielles sur allgemeinen Relativitatstheorte; von A. Einstein

Eine Reihe von Publikationen der letzten Zeit, insbesondere e neultch in diesen Annales 28. Heft 18 erschinnen scharfanige Arbeit von Kretschmann, veranlassen mich, nochals auf die Grundlagen der allgemeinen Relativitätthoorie röckrukomme. Dabei ist emein Zeit, lödigich die Grunddanken herauszuheben, wobei ich die Theorie als bekardt raussetze.

Die Theorie, wie ein ein heute vorschwebt, berybt auf drei uptgesichtspunkten, die allerdings keineswegt voneinander abhängig sind. Sie esten im folgenden kurs angeführt und arakterisiert und hierauf im nachfolgenden von einigen Seiten leonbet:

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b) Aquivalenzpriszig: Trägheit und Schwere und wessensich. Hierzus und ass den Ergebnissen der spreidlen Reivitästheorie folgt notwendig daß der symmetrische "Panmentaltensor" (q.,) die metrischen Eigenschaften des Raumes, a Trägheitzwerhalten der Körper in ihm, swise die Gravitationsrkungen bestimmt. Des darch des Pundamentaltensor betriebenen Raumstand wollen wir als "G-Fäld" bezeichnes.

c) Machaches Prinzip¹): Das G-Feld ist restlos durch die issen der Körper bestimmt. Da Masse und Energie nach

¹⁾ Bisher habe ich die Prinzipe a) und e) alcht auseinandergehatten, a hoer verwirrend wirkte. Den Namen "Machashes Prinzip" habe i denhalb geruhlt, weid dies Prinzip eise Veraligemeinerung der Machien Porderung bedautet, daß die Trägheit auf eine Wecheelwirtung : Körper auröckgeführt werden mösse.

918.

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 pre-general-relativistic theories can also be written in terms of coordinates that are reading on rods and clocks

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Die Theorie, wie ein ein heute vorschwebt, berybt auf drei uptgesichtspunkten, die allerdings keineswegt voneinander abhangig sind. Sie esten im folgenden kurs angeführt und arakterisiert und hierauf im nachfolgenden von einigen Seiten leonbet:

 a) Relativitätsprinzip: Die Naturgesetze sind nur Aussagen er zeiträumliche Koinzidenzen; sie finden deshalb ihren einzig tärlichen Ausdruck in allgemein kovarianten Gleichungen.

b) dyniadexprissip: Trägheit und Schwere und wesantich. Hirrsus und ass des Ergebnisses der spreidlen Rezvitätstheorie folgt notwendig, daß der symmetrische "Panmentältensor" (g.) die metrisches Eigenschaften des Raumes, a Trägheitverschlach der Körper in ihm, swie die Gravitationsrkungen bestimmt. Des darch des Pundamentaltensor betriebenen Raumzstaht welles wir als "G-Fäld" bezeichnes.

c) Machaches Prinzip¹): Das G-Feld ist restlos durch die assen der Körper bestimmt. Da Masse und Energie nach pre-general-relativistic theories contain not just statements about
coincidences but also statements about coordinate systems which serve for their description

 general relativity entails nothing more than relations between coincidences, the statements of which are independent of the choice of coordinates.

bisher habe ich die Prinzipe a) und e) nicht auseinandergehalten, a aber verwirrend wirkte. Den Namen "Machsebes Prinzip" habe ichabaib gewählt, weil dies Prinzip eine Veraligemeinerung der Machien Porderung beduetet, daß die Trägheit auf eine Wechselwirkung Körper zurrückgefühlt werden mösse.

Coordinates are not an adequate means of identifying world points physically. A physically meaningful statement about events is necessarily a statement about coincidences; that is, the prototype statement should be : 'Events *A*, *B*, *C*, ... took place at the same world point'*

"

- \blacksquare the value of the $g^{ij}(\boldsymbol{x}^i)$ is not an observable
- the value $A^i(x^i)$ at a point is not an observable
- the coincidence $g^{ij}(A^i)$ is an observables (in principle can be measured)

^{*}Bergmann and Komar, 1962, see also Bergmann, 1961, 1962.

Conclusion



What is left of the worldpoint? [...] It is not inconceivable that in the present formulation of general relativity the world point continues to exist as a relic from previous physical theories, to be discarded at the next stage of theoretical development ...^a

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^aBergmann, *The Riddle of Gravitation*, 1968.

Thanks!

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